

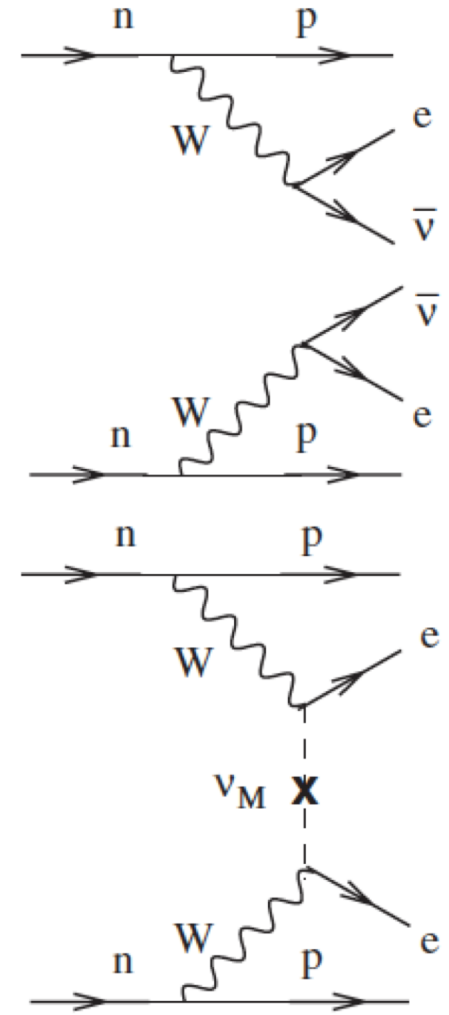
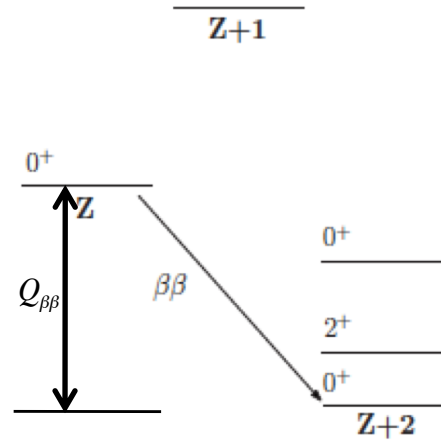
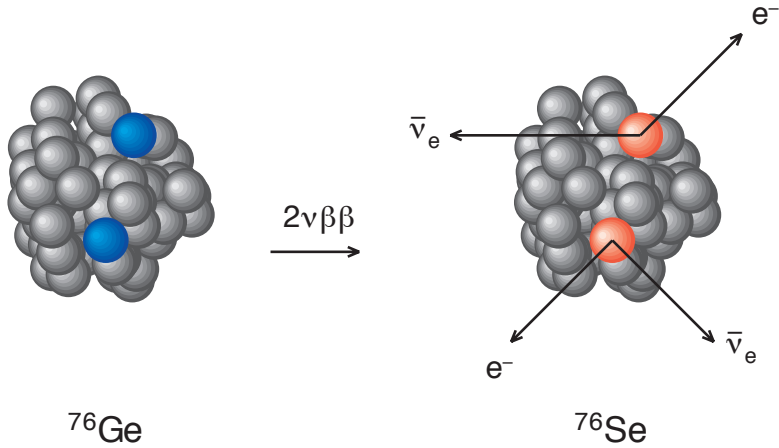
# Neutrinoless Double Beta Decay of Atomic Nuclei

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➤ Support from DOE grant DE-SC0015376 is acknowledged

# Classical Double Beta Decay Problem



$$\langle m_{\beta\beta} \rangle = \left| \sum_k m_k U_{ek}^2 \right|$$

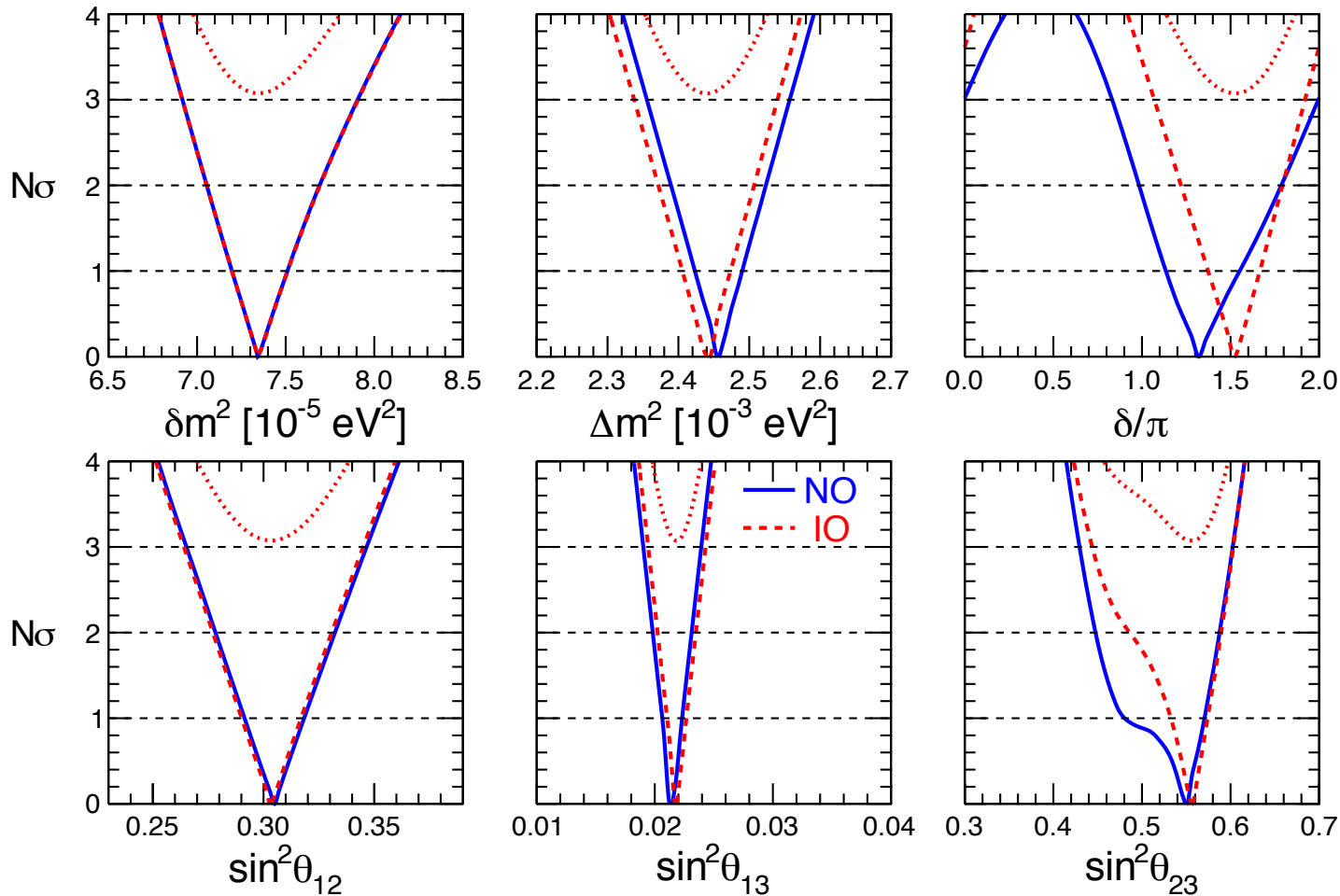
$$T_{1/2}^{-1}(0\nu) = G^{0\nu} (Q_{\beta\beta}) \left[ M^{0\nu}(0^+) \right]^2 \left( \frac{\langle m_{\beta\beta} \rangle}{m_e} \right)^2$$

$$\nu_e(x) = \sum_i U_{ei} \nu_i(x)$$

$$|\nu_e\rangle = \sum_i U_{ei}^* |\nu_i\rangle$$

# Neutrino oscillations parameters

LBL Acc + Solar + KamLAND + SBL Reactors + Atmos



Bari group:

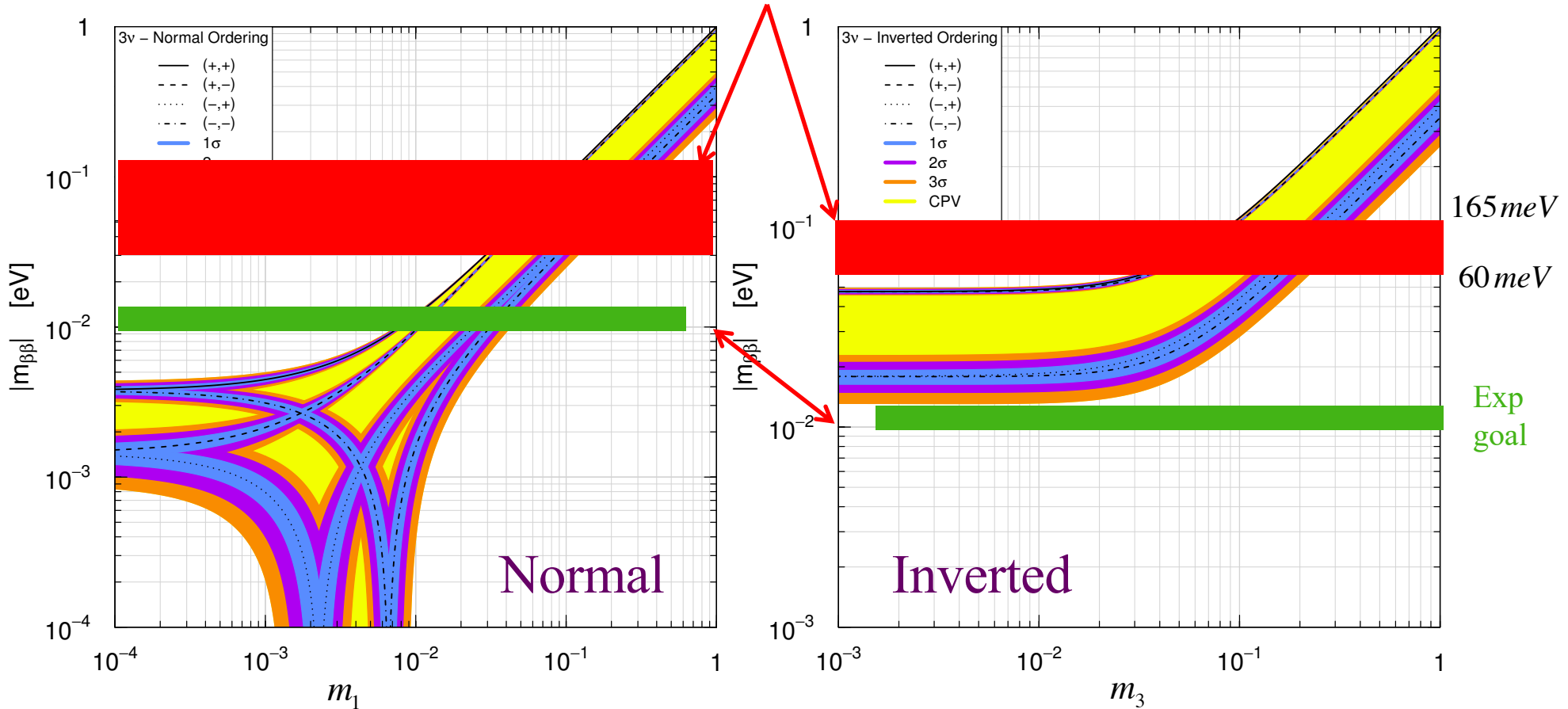
[arxiv.org/1804.09678](https://arxiv.org/1804.09678)  
 Prog. Part. Nucl. Phys.  
 102, 48 (2018)

1806.11051 review: Normal ordering favored at  $3.5\sigma$  !!

# Neutrino $\beta\beta$ effective mass

arxiv:1507.08204

KamLAND – Zen, PRL 117, 082503 (2016):  $^{136}\text{Xe}$



$$|m_{\beta\beta}| = \left| \sum_{k=1}^3 m_k U_{ek}^2 \right| = \left| c_{12}^2 c_{13}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right|$$

$$\phi_2 = \alpha_2 - \alpha_1 \quad \phi_3 = -\alpha_1 - 2\delta$$

$$\Leftrightarrow T_{1/2}^{-1}(0\nu) = G^{0\nu} (Q_{\beta\beta}) \left[ M^{0\nu}(0^+) \right]^2 (\eta_{0\nu})^2$$

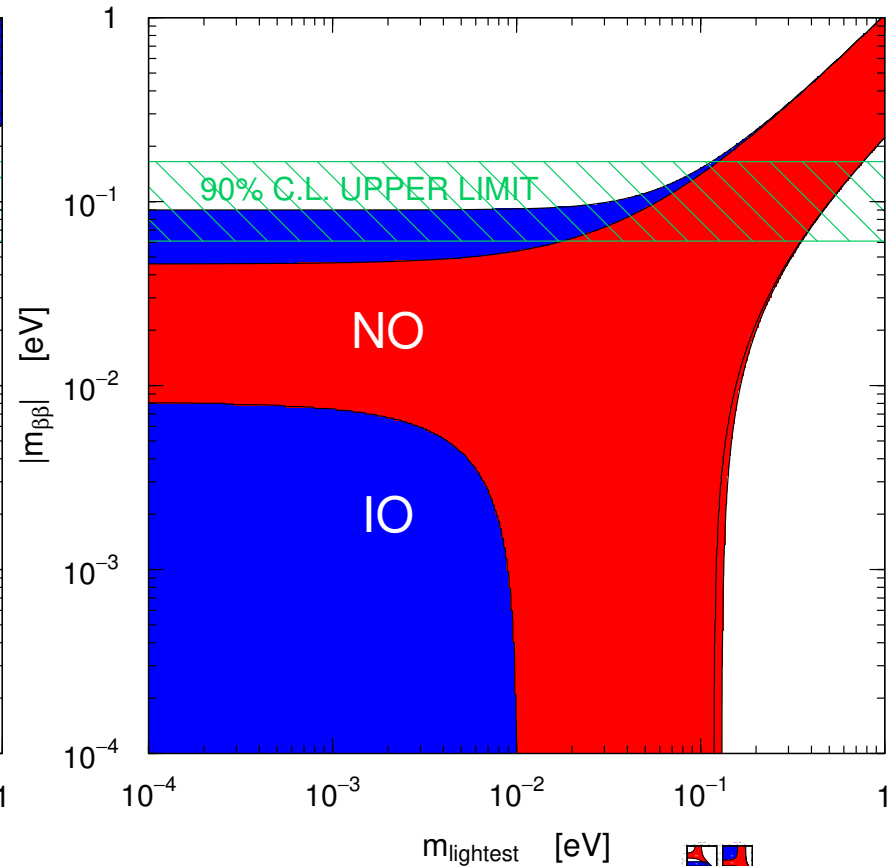
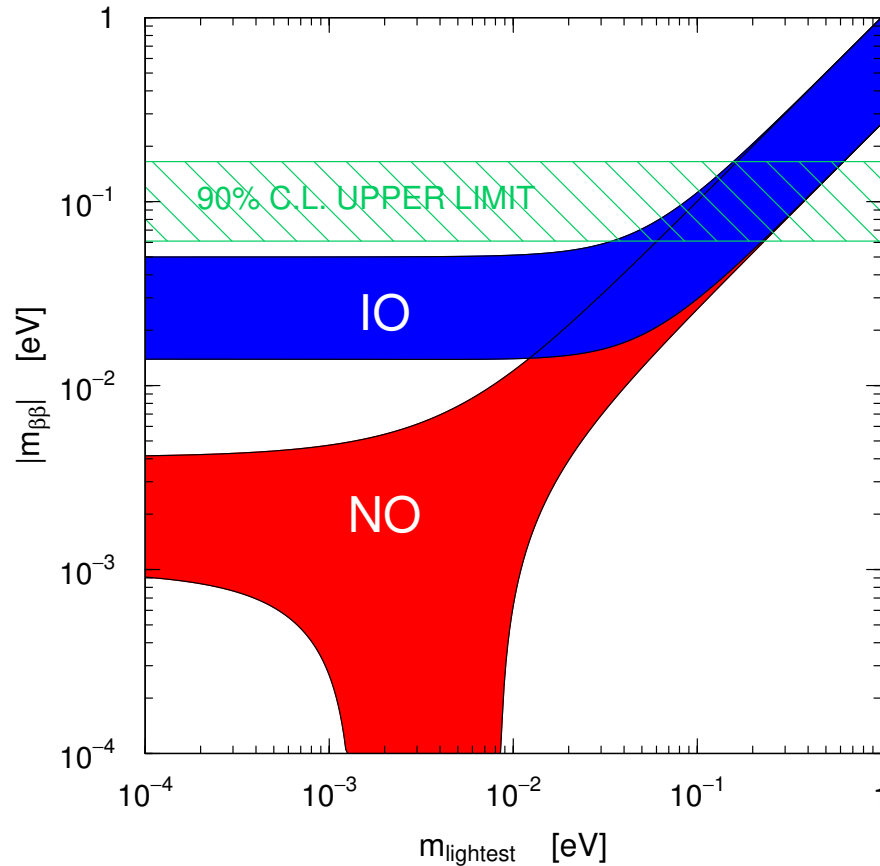
$$\eta_{0\nu} = \frac{|m_{\beta\beta}|}{m_e}$$



# $0\nu\beta\beta$ decay mass mechanism

3 neutrino flavors

3+1(sterile) neutrino flavors



$$|m_{\beta\beta}| = \left| \sum_{k=1}^3 m_k U_{ek}^2 \right| = \left| c_{12}^2 c_{13}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right|$$

$$\phi_2 = \alpha_2 - \alpha_1 \quad \phi_3 = -\alpha_1 - 2\delta$$

$$\Leftrightarrow T_{1/2}^{-1}(0\nu) = G^{0\nu} (Q_{\beta\beta}) [M^{0\nu}(0^+)]^2 (\eta_{0\nu})^2$$

$$\eta_{0\nu} = \frac{|m_{\beta\beta}|}{m_e}$$

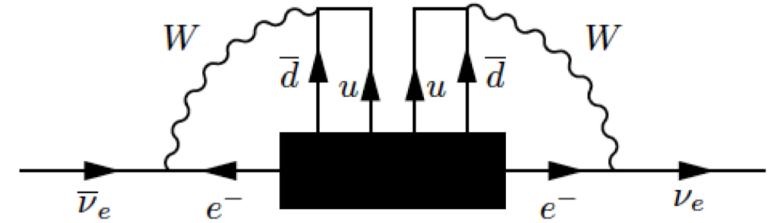
# The Black Box Theorems

## Black box I (electron neutrino)

J. Schechter and J.W.F Valle, PRD 25, 2951 (1982)

E. Takasugi, PLB 149, 372 (1984)

J.F. Nieves, PLB 145, 375 (1984)



$0\nu\beta\beta$  observed  $\Leftrightarrow$   
at some level

(i) Lepton number conservation is violated by 2 units.

(ii) Electron neutrinos are Majorana fermions (with  $m > 0$ ).

However:

M. Duerr et al, JHEP 06 (2011) 91

$$(\delta m_{\nu_e})_{BB} \sim 10^{-24} eV \ll \sqrt{|\Delta m_{32}^2|} \approx 0.05 eV$$

## Black box II (all flavors + oscillations)

M. Hirsch, S. Kovalenko, I. Schmidt, PLB 646, 106 (2006)

$0\nu\beta\beta$  observed  $\Leftrightarrow$   
at some level

(i) Lepton number conservation is violated by 2 units.

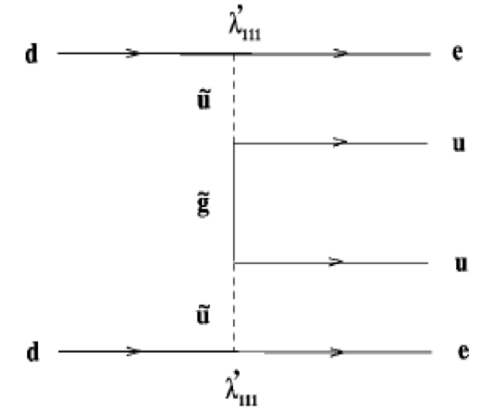
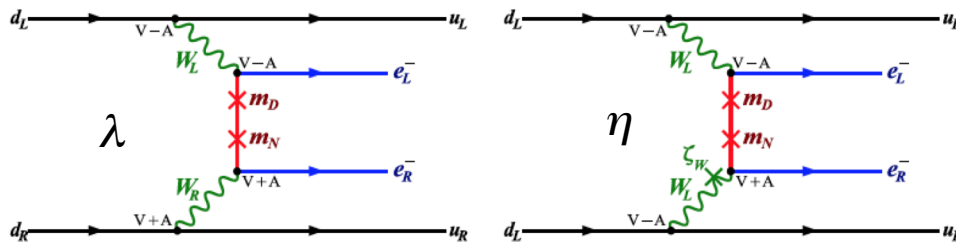
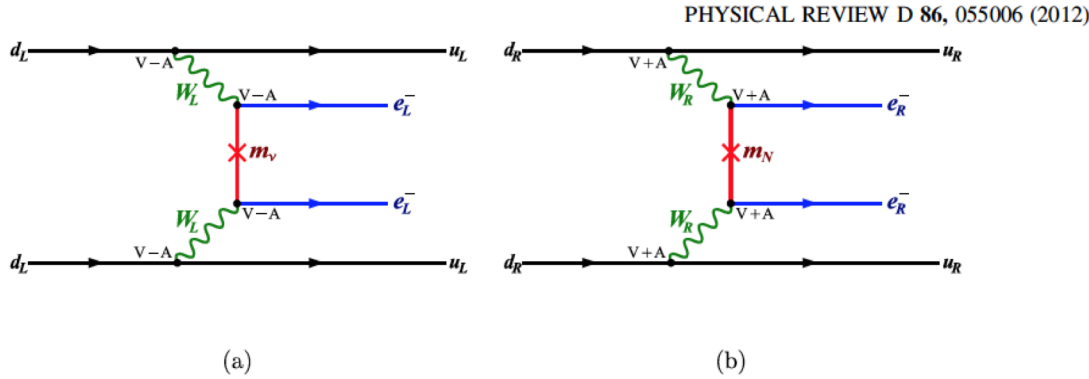
(ii) Neutrinos are Majorana fermions.

Regardless of the dominant  $0\nu\beta\beta$  mechanism!

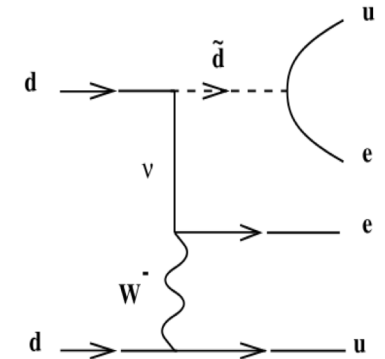
$$(iii) \quad \langle m_{\beta\beta} \rangle = \left| \sum_{k=1}^3 m_k U_{ek}^2 \right| = \left| c_{12}^2 c_{13}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right| > 0$$

# Other models: Left-Right symmetric model and SUSY R-parity violation

DAS *et al.*



Gluino exchange



Squark exchange

$$\left[ T_{1/2}^{0\nu} \right]^{-1} = G_{01} g_A^4 \left| \eta_{0\nu} M_{0\nu} + (\eta_{N_R}^L + \eta_{N_R}^R) M_{0N} + \eta_{\tilde{q}} M_{\tilde{q}} + \eta_{\lambda'} M_{\lambda'} + \eta_{\lambda} X_{\lambda} + \eta_{\eta} X_{\eta} \right|^2.$$

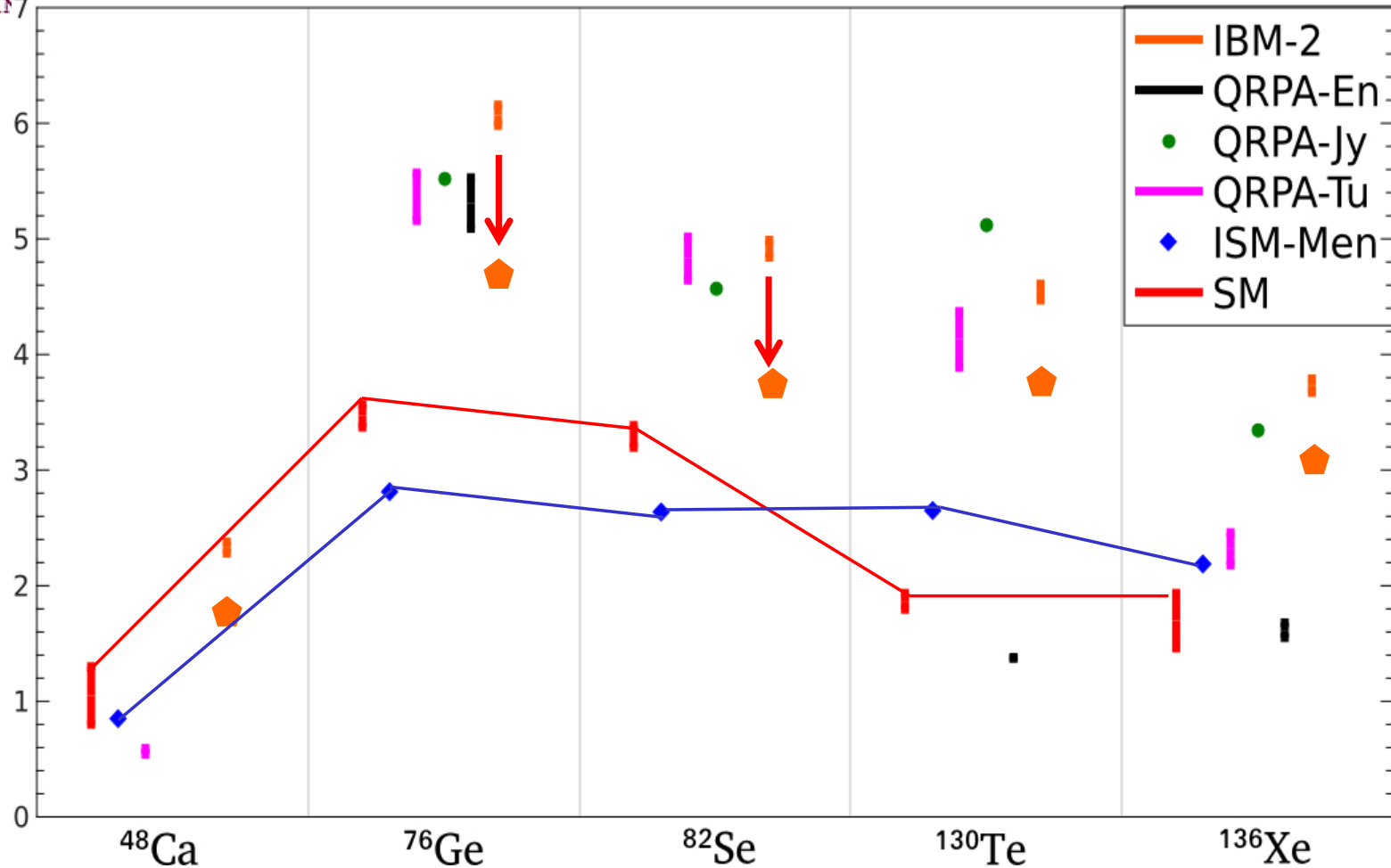
(e)

M. Horoi, A. Neacsu, PRD 93, 113014 (2016)

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# NME for the light-neutrino exchange mechanism



**IBA-2** J. Barea, J. Kotila, and F. Iachello, Phys. Rev. C **87**, 014315 (2013). **→** **IBM-2** PRC **91**, 034304 (2015)

**QRPA-En** M. T. Mustonen and J. Engel, Phys. Rev. C **87**, 064302 (2013).

**QRPA-Jy** J. Suhonen, O. Civitarese, Phys. NPA **847** 207–232 (2010).

**QRPA-Tu** A. Faessler, M. Gonzalez, S. Kovalenko, and F. Simkovic, arXiv:1408.6077

**ISM-Men** J. Menéndez, A. Poves, E. Caurier, F. Nowacki, NPA **818** 139–151 (2009).

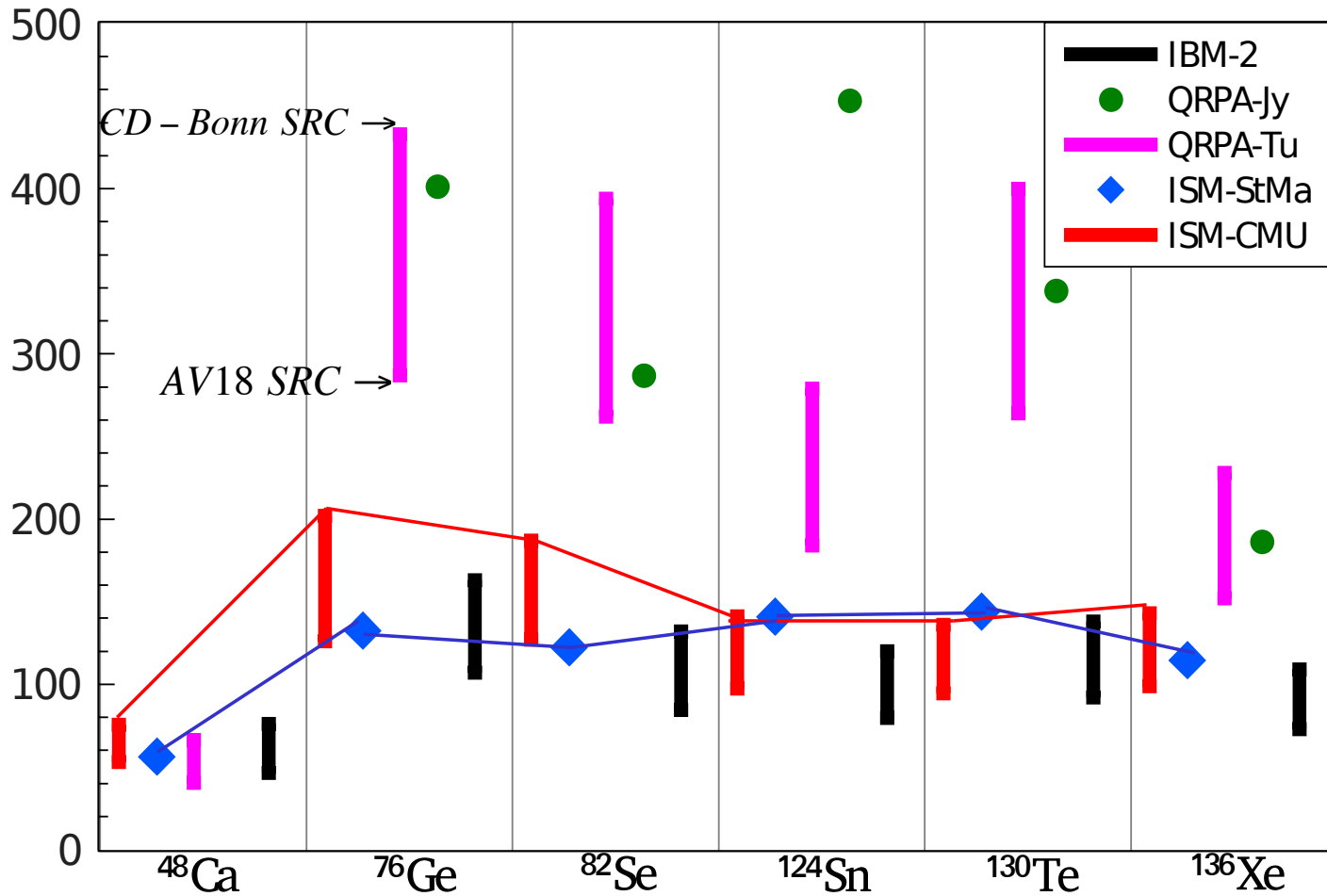
**SM** M. Horoi et. al. PRC **88**, 064312 (2013), PRC **89**, 045502 (2014), PRC **89**, 054304 (2014), PRC **90**, 051301(R) (2014), PRC **91**, 024309 (2015), PRL **110**, 222502 (2013), PRL **113**, 262501(2014).

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# Heavy neutrino-exchange NME

$M_{0N}$



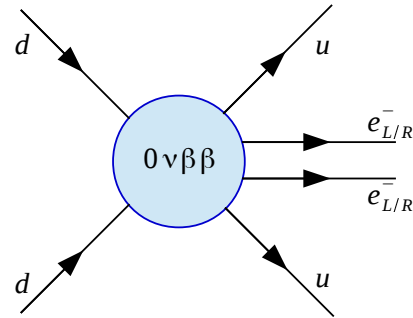
**IBA-2** J. Barea, J. Kotila, and F. Iachello, Phys. Rev. C **87**, 014315 (2013).

**QRPA-Tu** A. Faessler, M. Gonzalez, S. Kovalenko, and F. Simkovic, arXiv:1408.6077.

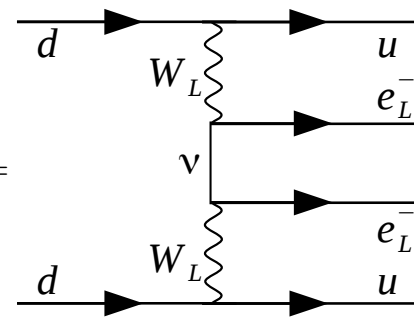
**QRPA-Jy** J. Hivarynen and J. Suhonen, PRC 91, 024613 (2015), **ISM-StMa** J. Menendez, private communication.

**ISM-CMU** M. Horoi et. al. PRC **88**, 064312 (2013), PRC **90**, PRC **89**, 054304 (2014), PRC **91**, 024309 (2015), PRL **110**, 222502 (2013).

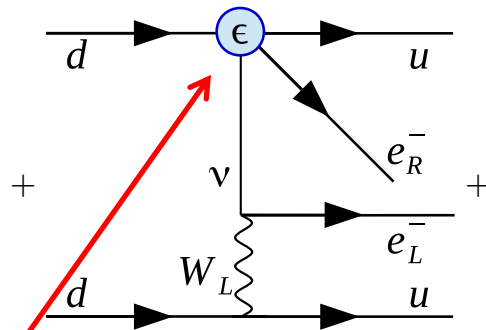
# Effective field theory approach



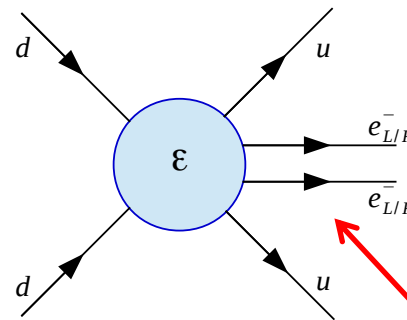
(a) The generic  $0\nu\beta\beta$  decay diagram at the quark-level.



(b) Light left-handed neutrino exchange diagram.



(c) The long-range part of the  $0\nu\beta\beta$  diagram.



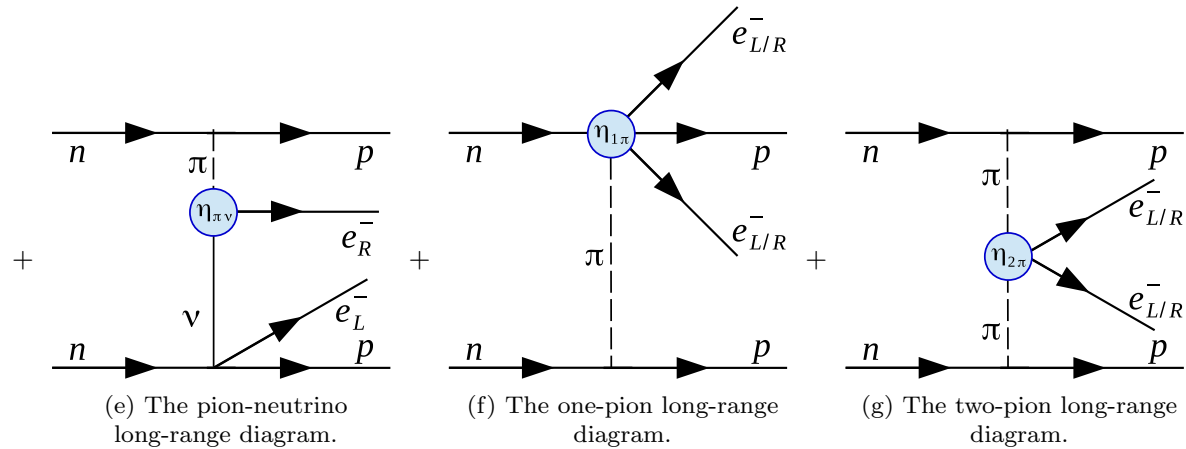
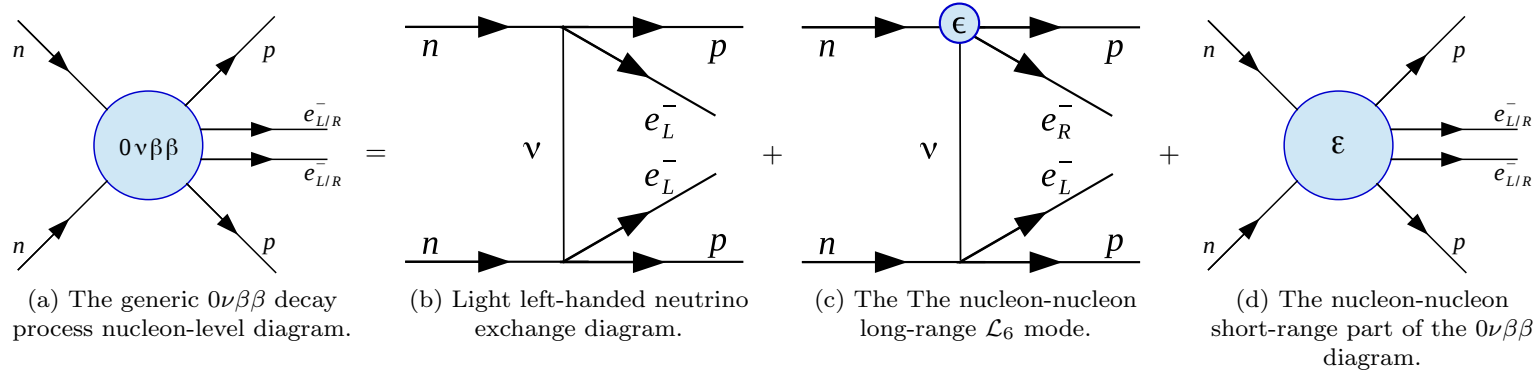
(d) The short-range part of the  $0\nu\beta\beta$  diagram.

arxiv:1706.05391

$$\mathcal{L}_6 = \frac{G_F}{\sqrt{2}} \left[ j_{V-A}^\mu J_{V-A,\mu}^\dagger + \sum_{\alpha,\beta}^* \epsilon_\alpha^\beta j_\beta J_\alpha^\dagger \right]$$

$$\mathcal{L}_9 = \frac{G_F^2}{2m_p} \left[ \epsilon_1 J J j + \epsilon_2 J^{\mu\nu} J_{\mu\nu} j + \epsilon_3 J^\mu J_\mu j \right. \\ \left. + \epsilon_4 J^\mu J_{\mu\nu} j^\nu + \epsilon_5 J^\mu J j_\mu \right],$$

# Effective field theory after hadronization



$$\left[ T_{1/2}^{0\nu} \right]^{-1} = g_A^4 \left[ \sum_i |\mathcal{E}_i|^2 \mathcal{M}_i^2 + \text{Re} \left[ \sum_{i \neq j} \mathcal{E}_i \mathcal{E}_j \mathcal{M}_{ij} \right] \right]$$

$$\mathcal{E}_{2-7} = \{ \epsilon_{V-A}^{V+A}, \epsilon_{V+A}^{V+A}, \epsilon_{S \pm P}^{S+P}, \epsilon_{TL}^{TR}, \epsilon_{TR}^{TR}, \eta_{\pi\nu} \}$$

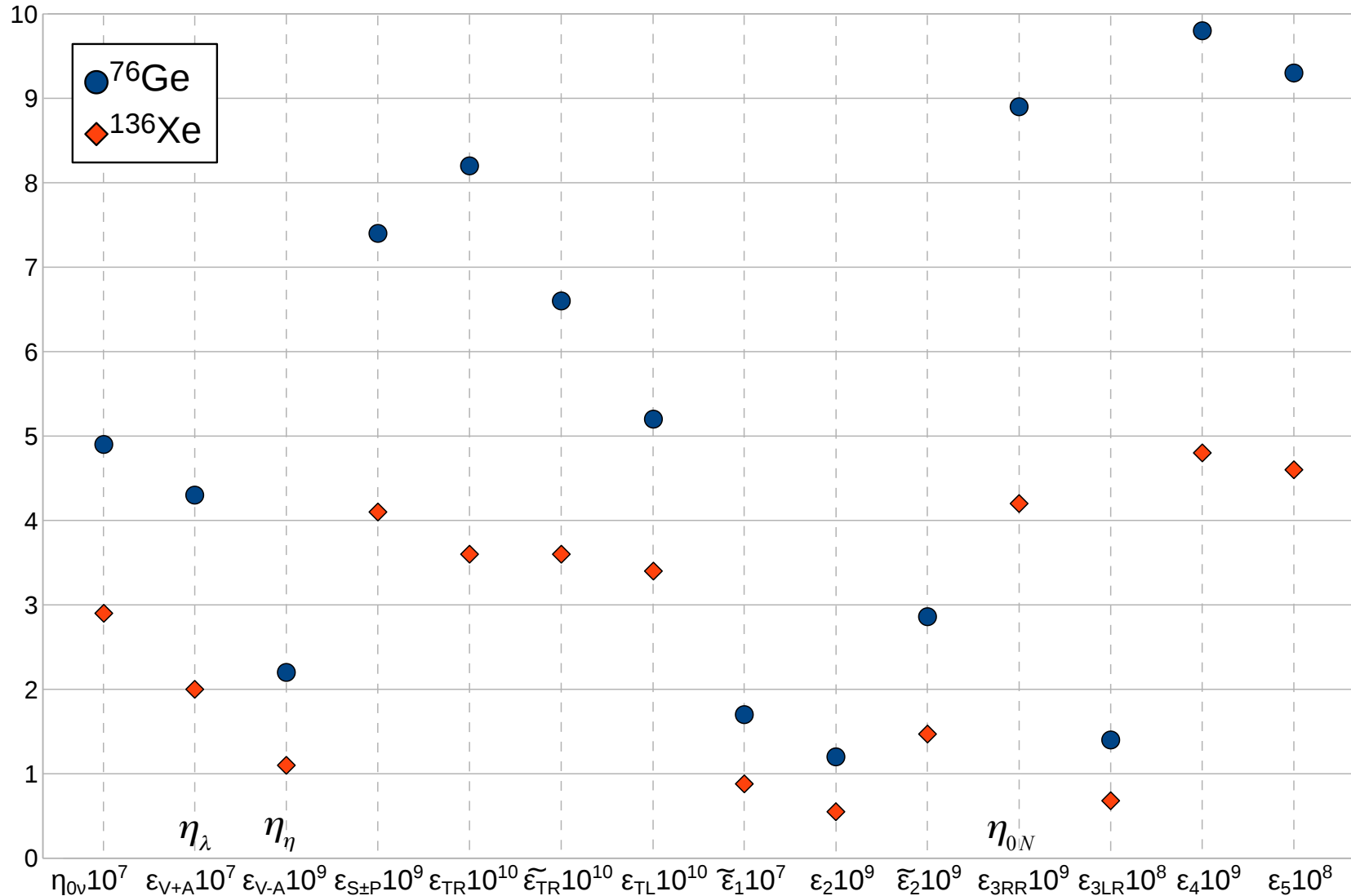
$$\mathcal{E}_{8-15} = \{ \epsilon_1, \epsilon_2, \epsilon_3^{LLz(RRz)}, \epsilon_3^{LRz(RLz)}, \epsilon_4, \epsilon_6, \eta_{1\pi}, \eta_{2\pi} \}$$

# One coupling dominance

$$[T_{1/2}^{0\nu}]^{-1} = g_A^4 \left[ \sum_i |\mathcal{E}_i|^2 \mathcal{M}_i^2 + \text{Re} \left[ \sum_{i \neq j} \mathcal{E}_{ij} \mathcal{M}_{ij} \right] \right]$$

$$\mathcal{E}_{2-7} = \{ \epsilon_{V-A}^{V+A}, \epsilon_{V+A}^{V+A}, \epsilon_{S\pm P}^{S+P}, \epsilon_{TL}^{TR}, \epsilon_{TR}^{\tilde{TR}}, \eta_{\pi\nu} \}$$

$$\mathcal{E}_{8-15} = \{ \epsilon_1, \epsilon_2, \epsilon_3^{LLz(RRz)}, \epsilon_3^{LRz(RLz)}, \epsilon_4, \epsilon_6, \eta_{1\pi}, \eta_{2\pi} \}$$



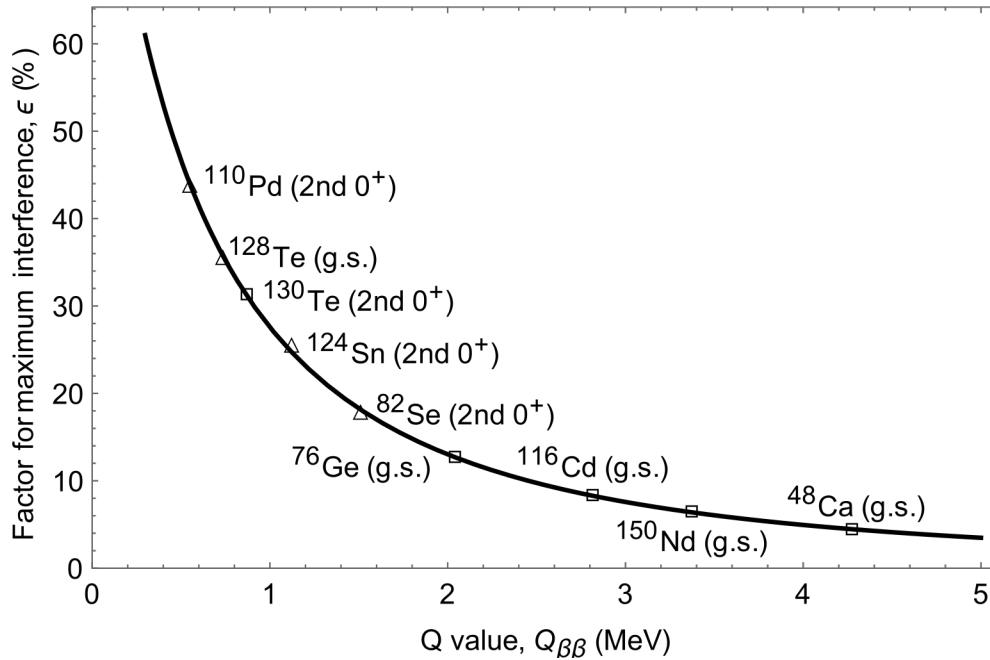
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$$T_{1/2}^{0\nu}({}^{76}\text{Ge}) > 5.3 \times 10^{25} \text{ years} \quad T_{1/2}^{0\nu}({}^{136}\text{Xe}) > 1.1 \times 10^{26} \text{ years}$$



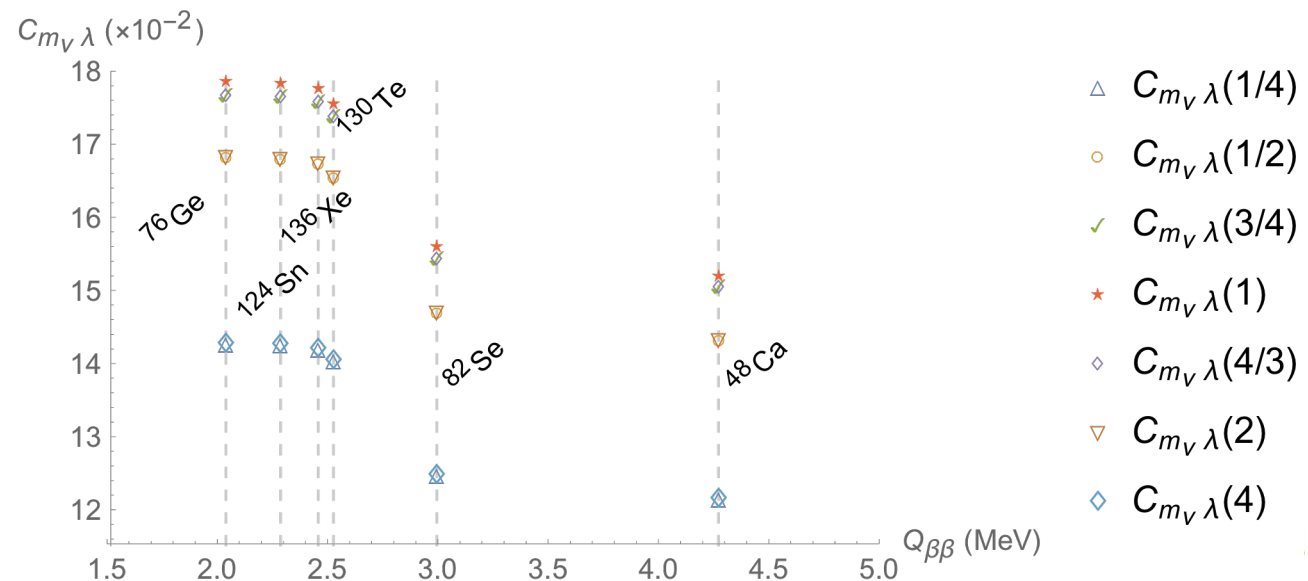
# Small interference effects



Interference between mass mechanism and heavy neutrino mechanism: F. Ahmed, A. Neacsu, and M. Horoi, Phys. Lett. B 769, 299 (2017).

Interference between mass mechanism and lambda mechanism: F. Ahmed, and M. Horoi, in preparation.

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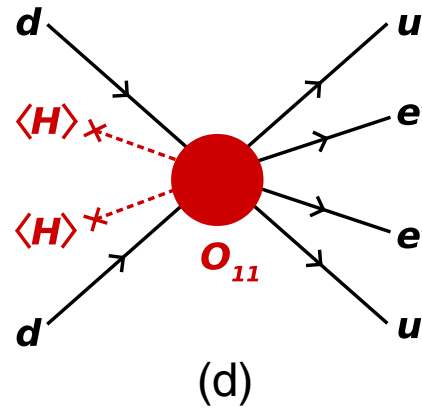
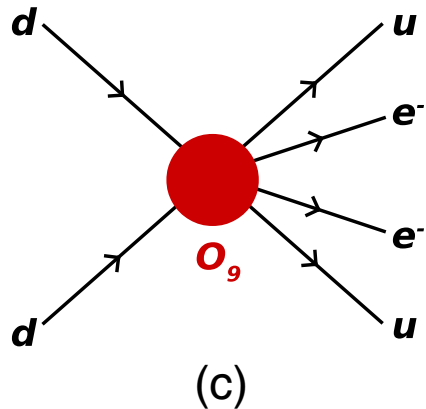
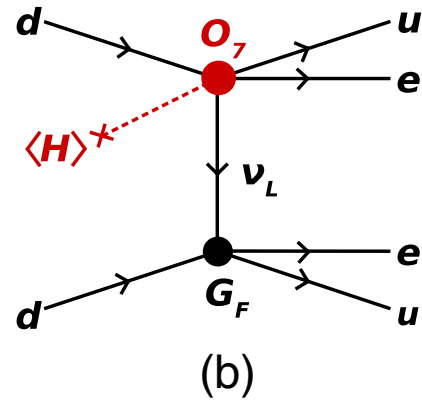
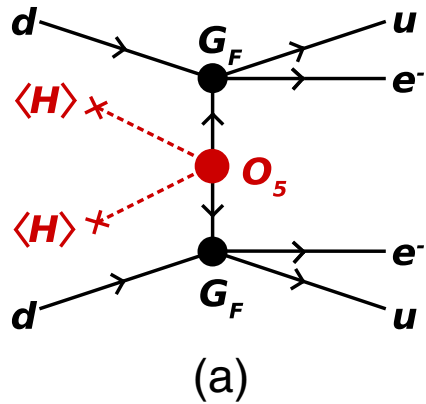


# Consequences: - scales for new physics

## - baryogenesis via leptogenesis

PHYSICAL REVIEW D **92**, 036005 (2015)

$$\mathcal{L}_D = \frac{g}{\Lambda_D^{D-4}} \mathcal{O}_D$$



$$m_e \bar{\epsilon}_5 = \frac{g^2 v^2}{\Lambda_5}, \quad \frac{G_F \bar{\epsilon}_7}{\sqrt{2}} = \frac{g^3 v}{2\Lambda_7^3},$$

$$\frac{G_F^2 \bar{\epsilon}_9}{2m_p} = \frac{g^4}{\Lambda_9^5}, \quad \frac{G_F^2 \bar{\epsilon}_{11}}{2m_p} = \frac{g^6 v^2}{\Lambda_{11}^7}$$

$g \approx 1 \quad v = 174 \text{ GeV}$  (Higgs expectation value)

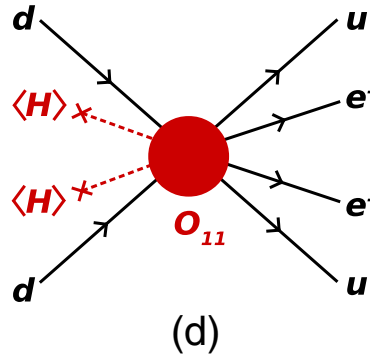
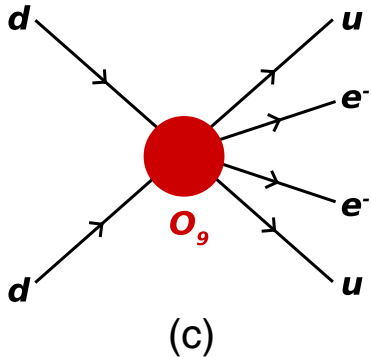
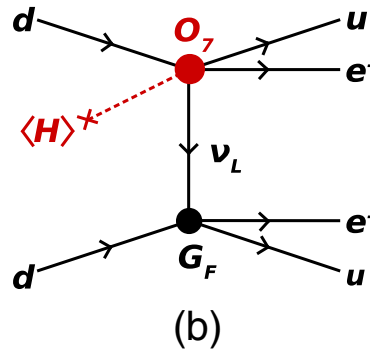
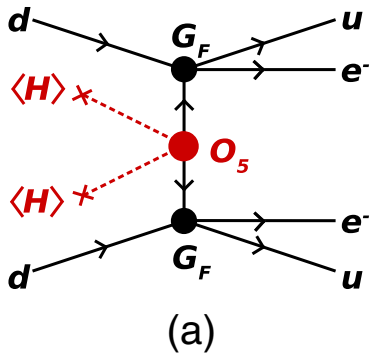
$\mathcal{O}_D$	$\bar{\epsilon}_D$	$\Lambda_D \text{ (GeV)}$
$\mathcal{O}_5$	$2.8 \times 10^{-7}$	$2.12 \times 10^{14}$
$\mathcal{O}_7$	$2.0 \times 10^{-7}$	$3.75 \times 10^4$
$\mathcal{O}_9$	$1.5 \times 10^{-7}$	$2.48 \times 10^3$
$\mathcal{O}_{11}$	$1.5 \times 10^{-7}$	$1.16 \times 10^3$

# Consequences: - scales for new physics

## - baryogenesis via leptogenesis

PHYSICAL REVIEW D **92**, 036005 (2015)

$$\mathcal{L}_D = \frac{g}{(\Lambda_D)^{D-4}} \mathcal{O}_D$$



$$m_e \bar{\epsilon}_5 = \frac{g^2 (yv)^2}{\Lambda_5}, \quad \frac{G_F \bar{\epsilon}_7}{\sqrt{2}} = \frac{g^3 (yv)}{2(\Lambda_7)^3},$$

$$\frac{G_F^2 \bar{\epsilon}_9}{2m_p} = \frac{g^4}{(\Lambda_9)^5}, \quad \frac{G_F^2 \bar{\epsilon}_{11}}{2m_p} = \frac{g^6 (yv)^2}{(\Lambda_{11})^7}$$

TABLE VIII. The BSM effective scale (in GeV) for different dimension-D operators at the present  $^{136}\text{Xe}$  half-life limit ( $\Lambda_D^0$ ) and for  $T_{1/2} \approx 1.1 \times 10^{28}$  years ( $\Lambda_D$ ).

$\mathcal{O}_D$	$\bar{\epsilon}_D$	$\Lambda_D^0(y=1)$	$\Lambda_D^0(y=y_e)$	$\Lambda_D(y=y_e)$
$\mathcal{O}_5$	$2.8 \cdot 10^{-7}$	$2.12 \cdot 10^{14}$	1904	19044
$\mathcal{O}_7$	$2.0 \cdot 10^{-7}$	$3.75 \cdot 10^4$	541	1165
$\mathcal{O}_9$	$1.5 \cdot 10^{-7}$	$2.47 \cdot 10^3$	2470	3915
$\mathcal{O}_{11}$	$1.5 \cdot 10^{-7}$	$1.16 \cdot 10^3$	31	43

$$\eta_N \propto \frac{1}{m_{W_R}^4 m_N}$$

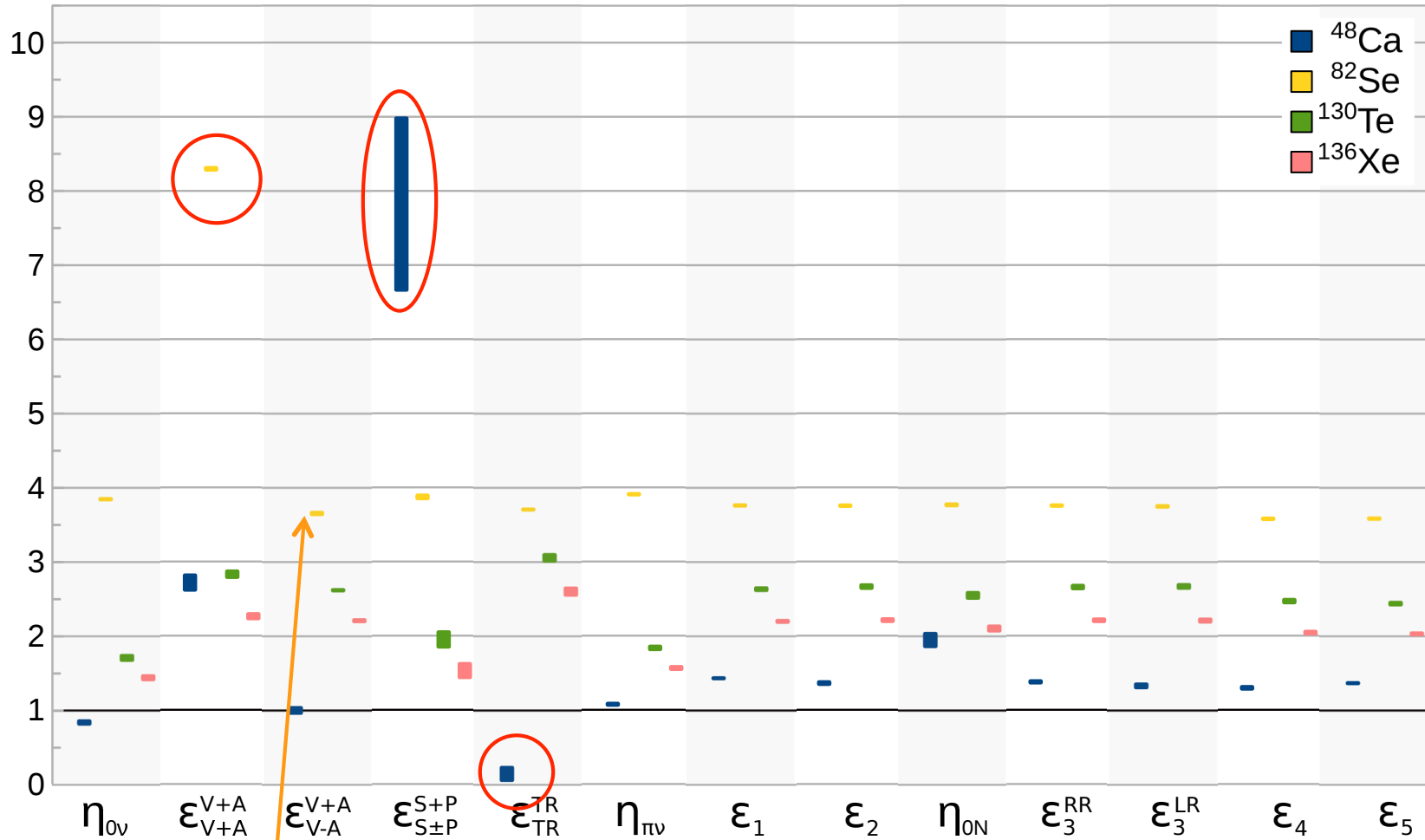
$$g \approx 1 \quad v = 174 \text{ GeV} \quad y_e = 3 \times 10^{-6} \text{ electron mass Yukawa}$$

One coupling dominance: which one?

$$[T_{1/2}^{0\nu}]^{-1} = g_A^4 \left[ \sum_i |\mathcal{E}_i|^2 \mathcal{M}_i^2 + \text{Re} \left[ \sum_{i \neq j} \mathcal{E}_i \mathcal{E}_j^* \mathcal{M}_{ij} \right] \right]$$

$T[{}^{76}\text{Ge}]/T[{}^A\text{Z}]$

CMU Hamiltonians



Super-NEMO

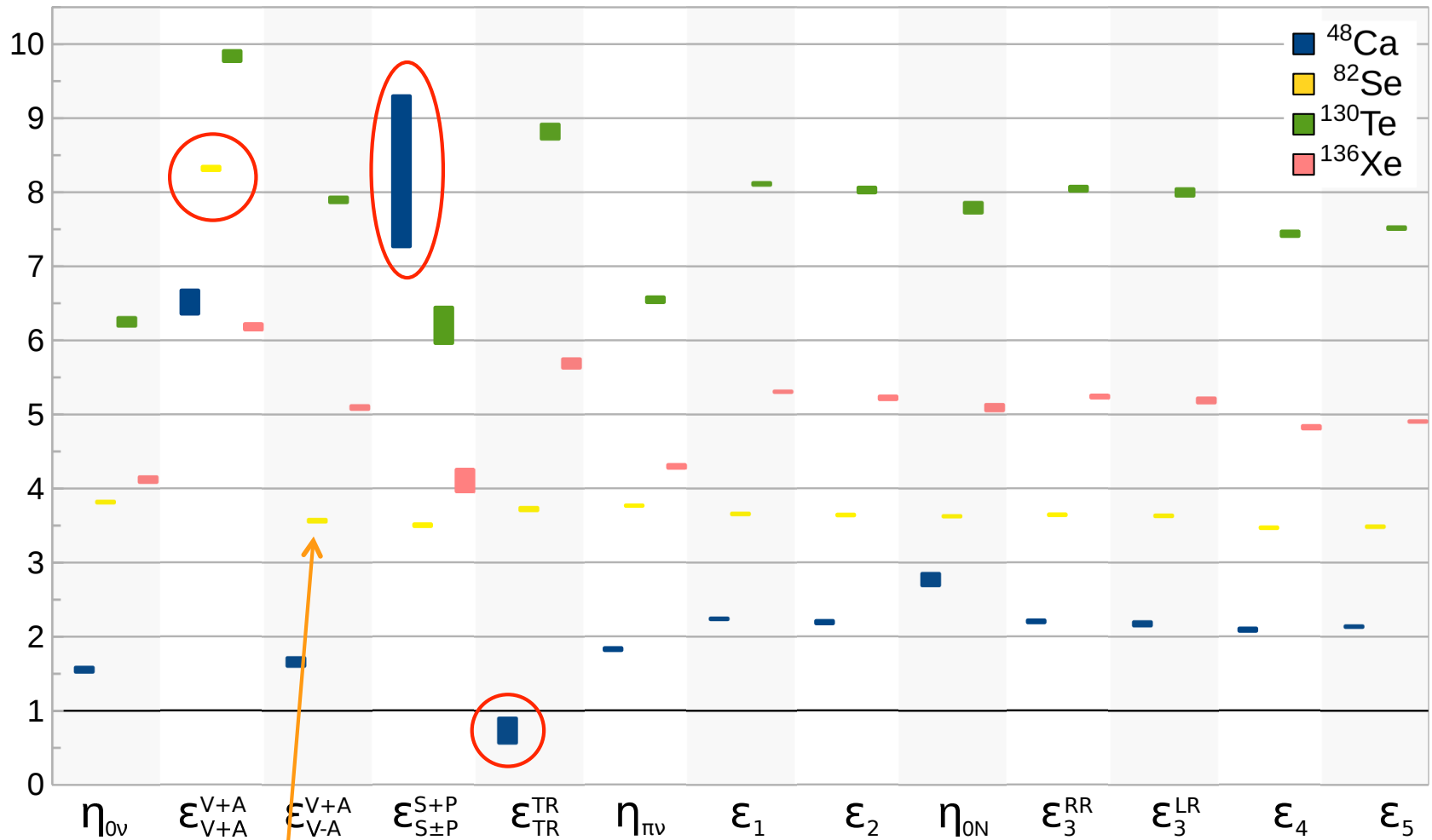
arxiv:1801.04496

One coupling dominance: which one?

$$[T_{1/2}^{0\nu}]^{-1} = g_A^4 \left[ \sum_i |\mathcal{E}_i|^2 \mathcal{M}_i^2 + \text{Re} \left[ \sum_{i \neq j} \mathcal{E}_{ij} \mathcal{M}_{ij} \right] \right]$$

T[<sup>76</sup>Ge]/T[AZ]

Strasbourg-Madrid Hamiltonians



Super-NEMO

## Neutrino Propagation in Nuclear Medium and Neutrinoless Double- $\beta$ Decay

S. Kovalenko,<sup>1</sup> M. I. Krivoruchenko,<sup>2,3</sup> and F. Šimkovic<sup>4,5,6</sup>

$$\mathcal{L}_{\text{eff}} = \frac{1}{\Lambda^2} \sum_{i,j,q} (g_{ij}^q \bar{\nu}_{Li}^C \nu_{Lj} \cdot \bar{q}q + \text{H.c.})$$

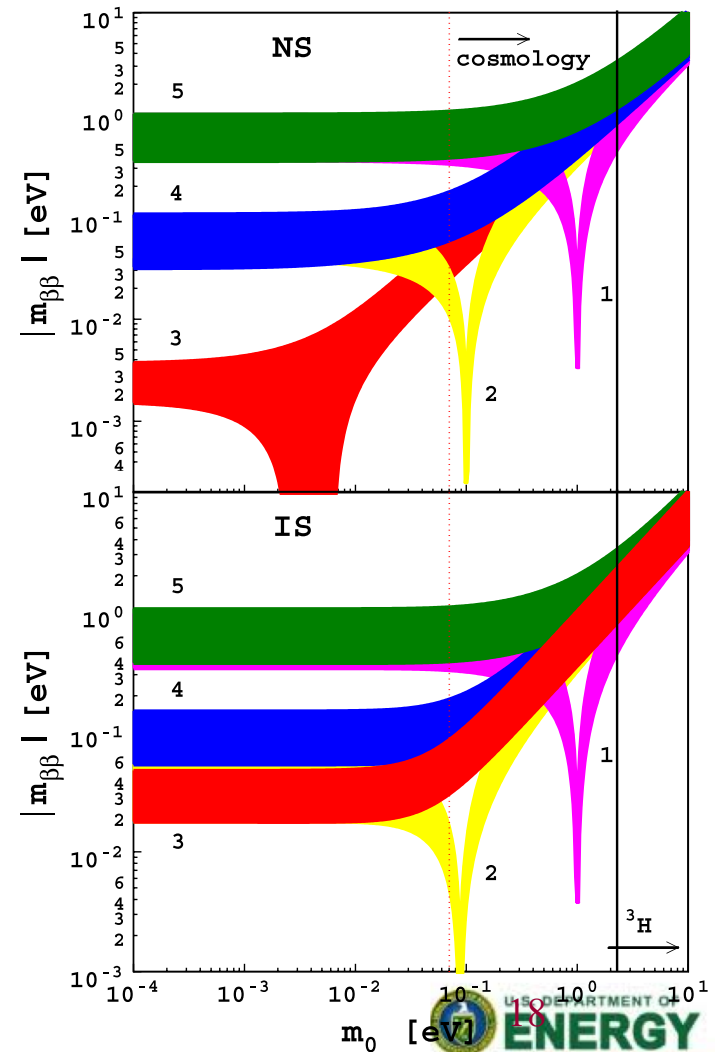
$$+ \frac{1}{\Lambda^3} \sum_{i,j,q} h_{ij}^q \bar{\nu}_{Li} i\gamma^\mu \overleftrightarrow{\partial}_\mu \nu_{Lj} \cdot \bar{q}q,$$

$$\mathcal{L}_{\text{eff}} = \frac{\langle \bar{q}q \rangle}{\Lambda_{\text{LNV}}^2} (\bar{\nu}_{Li}^C g_{ij} \nu_{Lj} + \text{H.c.})$$

$$+ \frac{\langle \bar{q}q \rangle}{\Lambda^3} \bar{\nu}_{Li} h_{ij} i\gamma^\mu \overleftrightarrow{\partial}_\mu \nu_{Lj},$$

$$m_{\beta\beta} = \sum_{i=1}^n (V_{ei}^L)^2 \xi_i \frac{|m_i - \langle \bar{q}q \rangle g|}{(1 - \langle \bar{q}q \rangle h)^2}.$$

How about some other contributions from SM? E.g. high density atomic electrons.



# Neutrinos in atomic nuclei

Atomic nucleus is a high electron density medium:

Consider 2 electrons in the lowest s-orbital of an  
Hydrogen-like atom

Si<sub>2</sub> dimer

*Electron density near nucleus:*

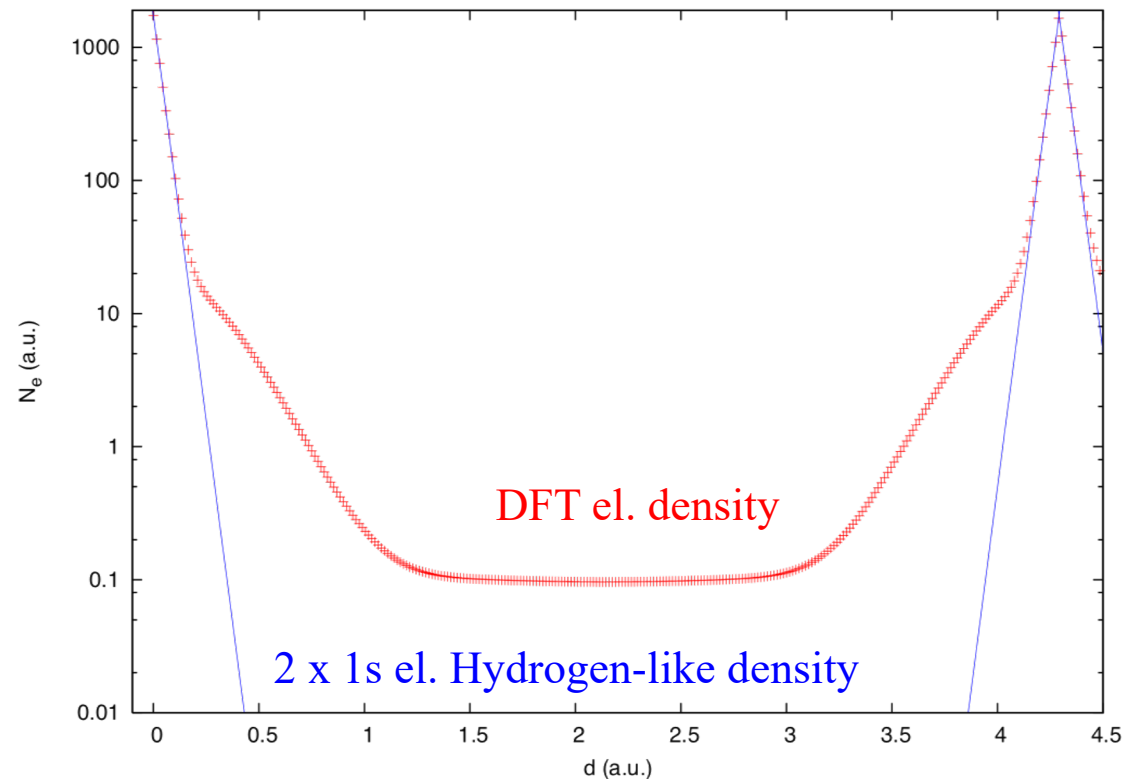
$$N_e(r) \approx \frac{2}{\pi} \left( \frac{Z}{a_B} \right)^3 e^{-2rZ/a_B}$$

*Electron density inside nucleus:*

$$N_e(0) \approx \frac{2}{\pi} \left( \frac{Z}{a_B} \right)^3$$

$$\rho_{Sun\ core} \approx 150 \text{ g / cm}^3$$

$$\text{Equivalent matter density: } \rho = m_N N_e = 1.67 \times 10^{-24} \frac{2}{\pi} \left( \frac{Z}{53} \right)^3 \text{ in g / cm}^3 \gg \rho_{Sun}$$

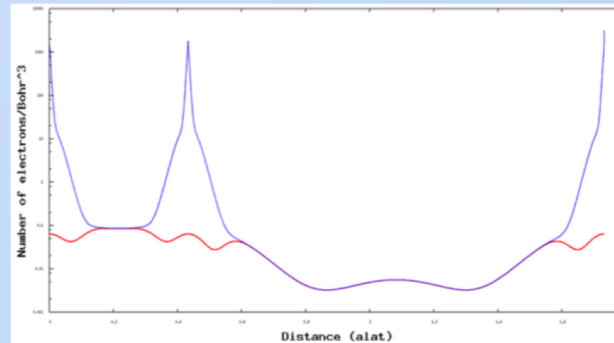
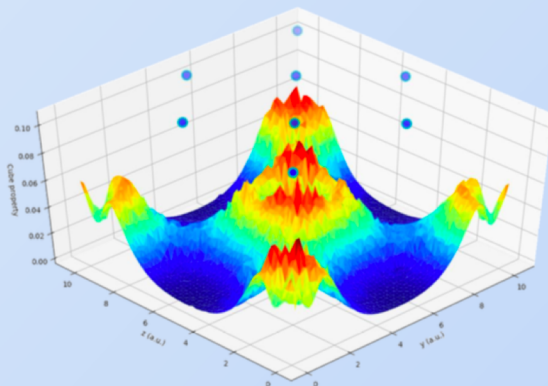


# Neutrinos in atomic nuclei

## Reconstructing Electron Charge Density T. E. Stearns, M. Fornari, M. Horoi, A. Zettel Dept. of Physics, Central Michigan University

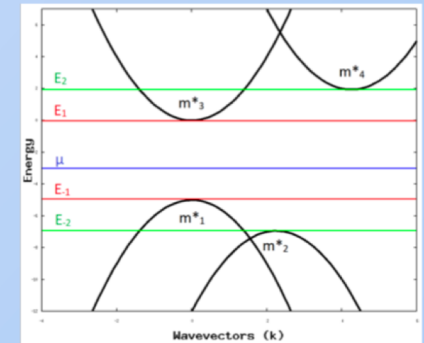
### Electronic Structure Collides with Neutrino Physics

A common method used by programs to compute the electronic structure involves treating the core electrons in an atom as “frozen”. This allows the calculation to focus on the electrons that participate in the chemical bonding, vastly improving computation speed. However, the resultant charge density is incomplete, since it contains only the valence electrons. Neutrinos may respond differently when the medium in which they travel is highly charged. In order to describe the interaction of neutrinos travelling through Earth, it is necessary to reconstruct the total charge density by adding the core electrons to the valence charge.



The total (blue) and valence (red) charge density of cubic silicon plotted logarithmically along a diagonal (1 1 1) line. Notice how both plots follow each other very closely, except in the areas where

- Charge Density is the probability that an electron will occupy a given volume of space.
- The program used during this project was *Quantum ESPRESSO*<sup>1</sup>, a plane wave based density functional theory calculation suite. It uses pseudopotentials to replace the core electrons.
- Pseudopotentials approximate the energy from the coulomb interaction between the core electrons and the valence electrons.
- Since the core electrons are chemically inactive, the density in the core regions of an atom in a material can be reconstructed from the pseudopotential.
- We tried a number of utilities to reconstruct the total density, including *Abinit*<sup>2</sup>, *GPAW*<sup>3</sup>, and *Critic2*<sup>4</sup>. *Quantum ESPRESSO* returned the best results.
- Eventually, this data will be used in a neutrino simulation by Adam Zettel and Dr. Horoi. Please see their poster for more information.



An example of parabolic band structure. The blue line is the chemical potential, or Fermi energy. The red and green lines



# Neutrinos in matter: local mass eigenstates

$$H = U_{vac} \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} U_{vac}^+ + \begin{pmatrix} 2EV & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

$$V(eV) = \pm 7.6 \times 10^{-14} m_p(g) N_e(cm^{-3})$$

$$H = U_{mat} \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta M_{21}^2 & 0 \\ 0 & 0 & \Delta M_{31}^2 \end{pmatrix} U_{mat}^+$$

(Anti)neutrinos are "emitted" in matter in the local (lowest) highest "mass eigenstates".

# Neutrinoless double beta decay in vacuum

$$A_{0\beta\beta} \propto NP = \langle 0 | T [\psi_{eL}(x_1) \psi_{eL}^T(x_2)] | 0 \rangle$$

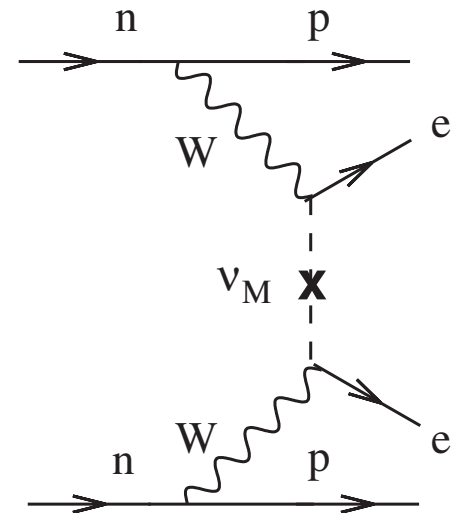
$$\psi_e(x) = \sum_{a=1}^{N(3)} U_{ea} \psi_a(x)$$

$$\begin{aligned} NP &= \sum_{a=1}^3 U_{ea}^2 \langle 0 | T [\psi_{aL}(x_1) \psi_{aL}^T(x_2)] | 0 \rangle \\ &= \sum_{a=1}^3 U_{ea}^2 \left[ -i \int \frac{d^4 p}{(2\pi)^4} \frac{m_a e^{-ip(x_1-x_2)}}{p^2 - m_a^2 + i\epsilon} P_L C \right] \end{aligned}$$

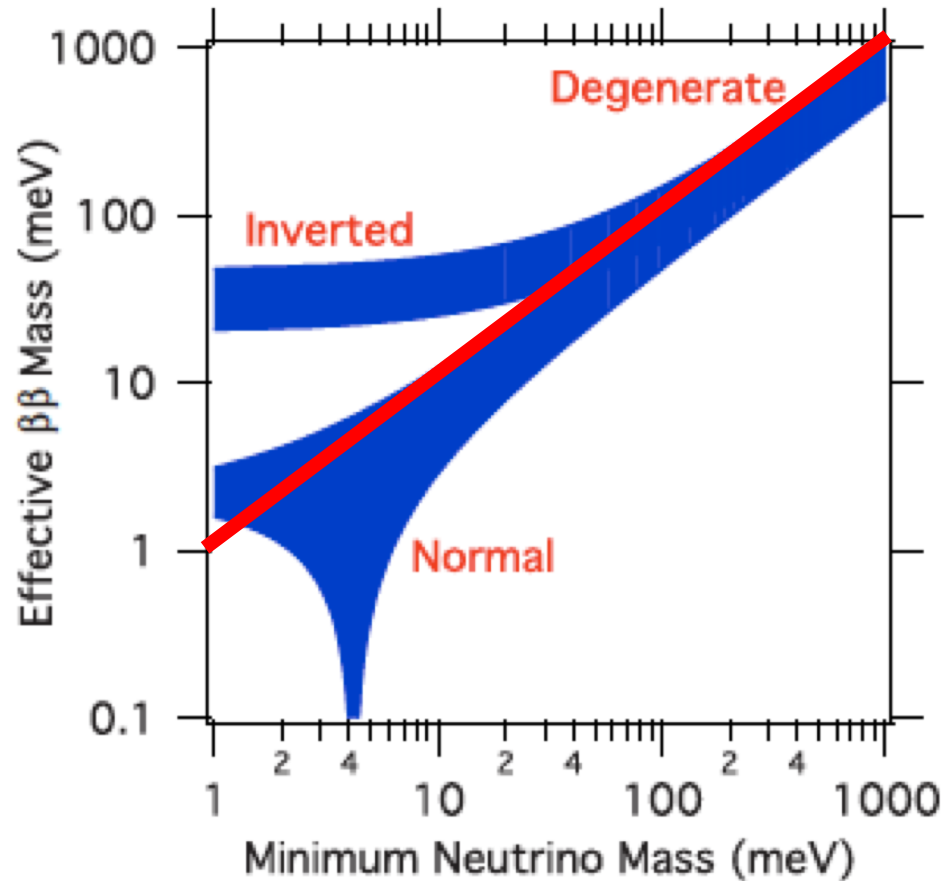
$$P_L = \frac{1}{2}(1 - \gamma^5) \quad \hat{\psi}(x) = C\psi^*(x)$$

$P_L C$  product is further used to process the electron current, and one finally gets:

$$\frac{1}{T_{1/2}} = G(Z, Q) |M_{0\nu}|^2 \left| \sum_{a=1}^3 U_{ea}^2 m_a \right|^2 / m_e^2$$



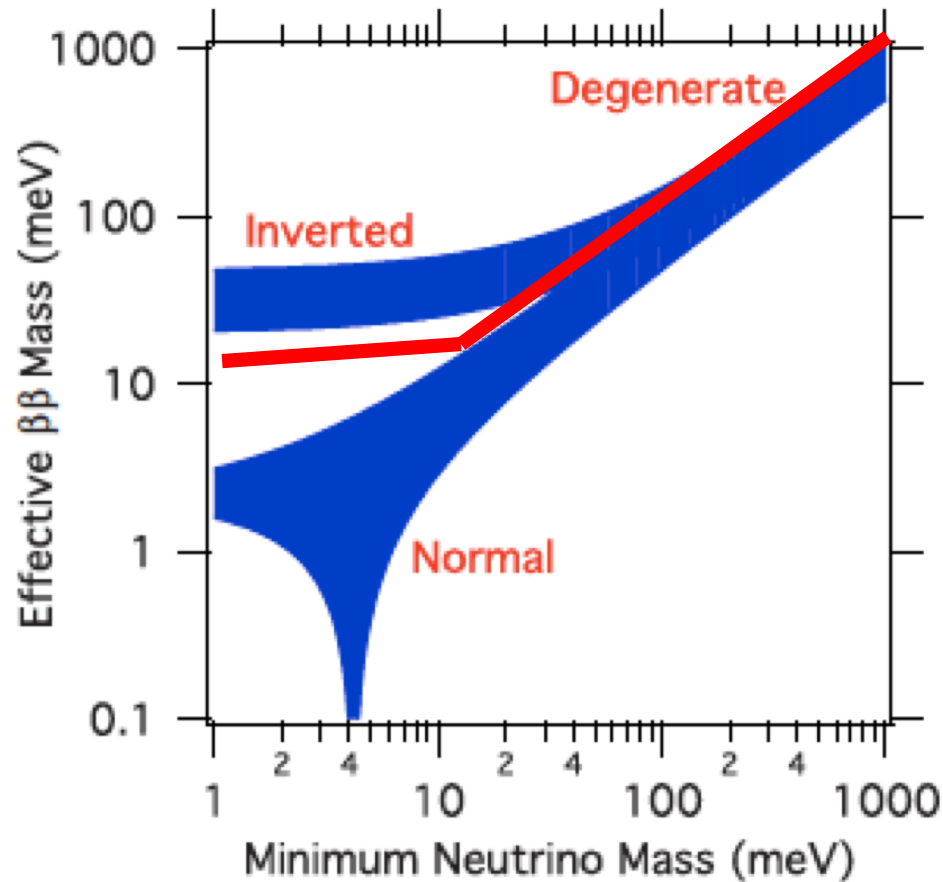
# Effective neutrino mass



$$\left| m_{\beta\beta} \right| = \left| U_{e j_i}^2 m_{j_i} \right| \quad \Leftarrow T_{1/2}^{-1}(0\nu) = G^{0\nu}(Q_{\beta\beta}) \left[ M^{0\nu}(0^+) \right]^2 (\eta_{0\nu})^2$$

$$\eta_{0\nu} = \frac{|m_{\beta\beta}|}{m_e}$$

# Effective neutrino mass



$$|m_{\beta\beta}| = \left| U_{e j_h}^2 m_{j_h} \right| \quad \Leftarrow T_{1/2}^{-1}(0\nu) = G^{0\nu} (Q_{\beta\beta}) [M^{0\nu}(0^+)]^2 (\eta_{0\nu})^2$$

$$\eta_{0\nu} = \frac{|m_{\beta\beta}|}{m_e}$$

# Neutrinoless double beta decay of atomic nuclei

Details are rather complex and can be found in arXiv:1803.06332

In short: the vacuum result stands!

$$\langle 0 | T \left[ \Phi_e^W(x_1) (\Phi_e^W(x_2))^T \right] | 0 \rangle = -i \sum_a U_{ea}^2 \int \frac{d^4 p}{(2\pi)^4} \frac{m_a e^{-ip(x_1-x_2)}}{p^2 - m_a^2 + i\epsilon} (i\sigma^2)$$

*In atomic nuclei NP = In vacuum NP*

*Vacuum result stands :*  $m_{\beta\beta} = \left| \sum_{a=1}^3 U_{ea}^2 m_a \right|$

$$P_L C = \begin{pmatrix} 0 & 0 \\ 0 & i\sigma^2 \end{pmatrix}$$

# Summary

- Neutrinoless DBD, if observed, will represent a big step forward in our understanding of the neutrinos, and of physics beyond the Standard Model.
- Ratios of half-lives for several isotopes are essential to account for alternative decay mechanisms.
- The effects of the high electron densities in atomic nuclei were investigated and they do not change the neutrino emission or detection, nor the  $0\nu\beta\beta$  outcome.
- These results look simple, but the road to them is complex. Consequences to matter effects neutrino oscillations could be interesting.